



Lorentz transformations via Pauli matrices

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Abstract.- We exhibit expressions, in terms of Pauli matrices, which directly generate Lorentz transformations in Minkowski space.

Key words.- Pauli matrices, Lorentz transformations, Infeld-van der Waerden symbols



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In space time an event is represented by $(x^j) = (ct, x, y, z)$, $j = 0, \dots, 3$, with the metric $(g_{jr}) = \text{Diag}(1, -1, -1, -1)$. If it is necessary to employ another frame of reference, then the new coordinates \tilde{x}^r are connected with x^j via the linear transformation:

$$\tilde{x}^j = L^j_r x^r \quad , \quad (1)$$

where the Lorentz matrix \underline{L} verifies the restriction :

$$L^j_a g_{rj} L^r_b = g_{ab} \quad , \quad (2)$$

because the Minkowskian line element must remain invariant under \underline{L} , that is, $\tilde{x}^r \tilde{x}_r = x^r x_r$.

From (2) we see that \underline{L} has six degrees of freedom, which permits to work with four complex numbers $\alpha, \beta, \gamma, \delta$ such that $\alpha\delta - \beta\gamma = 1$, then the components of homogeneous Lorentz transformation \underline{L} can be written in the form [1-4]:

$$\begin{aligned} L^0_0 &= \frac{1}{2}(\alpha\alpha^* + \beta\beta^* + \gamma\gamma^* + \delta\delta^*) \quad , & L^0_1 &= \frac{1}{2}(\alpha^*\beta + \gamma^*\delta) + \text{c.c.} \quad , \\ L^0_2 &= \frac{i}{2}(\alpha^*\beta + \gamma^*\delta) + \text{c.c.} \quad , & L^0_3 &= \frac{1}{2}(\alpha\alpha^* - \beta\beta^* + \gamma\gamma^* - \delta\delta^*) \quad , \\ L^1_0 &= \frac{1}{2}(\alpha^*\gamma + \beta^*\delta) + \text{c.c.} \quad , & L^1_1 &= \frac{1}{2}(\alpha^*\delta + \beta\gamma^*) + \text{c.c.} \quad , \\ L^1_2 &= \frac{i}{2}(\alpha^*\delta + \beta\gamma^*) + \text{c.c.} \quad , & L^1_3 &= \frac{1}{2}(\alpha^*\gamma - \beta^*\delta) + \text{c.c.} \quad , \\ L^2_0 &= \frac{i}{2}(\alpha\gamma^* - \beta^*\delta) + \text{c.c.} \quad , & L^2_1 &= \frac{i}{2}(\alpha\delta^* + \beta\gamma^*) + \text{c.c.} \quad , \\ L^2_2 &= \frac{1}{2}(\alpha^*\delta - \beta^*\gamma) + \text{c.c.} \quad , & L^2_3 &= \frac{i}{2}(\alpha\gamma^* + \beta^*\delta) + \text{c.c.} \quad , \\ L^3_0 &= \frac{1}{2}(\alpha\alpha^* + \beta\beta^* - \gamma\gamma^* - \delta\delta^*) \quad , & L^3_1 &= \frac{1}{2}(\alpha^*\beta - \gamma^*\delta) + \text{c.c.} \quad , \\ L^3_2 &= \frac{i}{2}(\alpha^*\beta - \gamma^*\delta) + \text{c.c.} \quad , & L^3_3 &= \frac{1}{2}(\alpha\alpha^* - \beta\beta^* - \gamma\gamma^* + \delta\delta^*) \quad , \end{aligned} \quad (3)$$

where c.c. means the complex conjugate of all the previous terms. Therefore, any complex 2x2 matrix [4-7]:

$$\underline{U} = \begin{pmatrix} \alpha & \beta \\ \gamma & \delta \end{pmatrix} \quad , \quad \text{Det } \underline{U} = \alpha\delta - \beta\gamma = 1 \quad , \quad (4)$$

generates a Lorentz matrix via (3).

The following relations, which are not explicitly in the literature, give us directly all the components (3):



$$L^{\mu}_{\nu} = -\frac{1}{2} U^{ar} \sigma^{\mu}_{aj} \sigma_{vbr} U^{\dagger bj} \quad , \quad \mu, \nu = 1, 2, 3$$

$$L^{\mu}_0 = \frac{1}{2} \sigma^{\mu}_{jr} Q^{jr} \quad , \quad \mu = 0, \dots, 3 \quad , \quad L^0_{\nu} = -\frac{1}{2} \sigma_{\nu jk} R^{jk} \quad , \quad \nu = 1, 2, 3 \quad (5)$$

such that:

$$U^{\dagger} = \begin{pmatrix} \alpha^* & \gamma^* \\ \beta^* & \delta^* \end{pmatrix} \quad , \quad Q = U U^{\dagger} \quad , \quad R = U^{\dagger} U \quad , \quad (6)$$

with the Infeld-van der Waerden symbols [8-11]:

$$\begin{aligned} (\sigma^0_{ab}) &= (\sigma_{0ab}) = \mathbb{I} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \quad , \quad (\sigma^1_{ab}) = (-\sigma_{1ab}) = \sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \\ (\sigma^2_{ab}) &= (-\sigma_{2ab}) = -\sigma_y = \begin{pmatrix} 0 & i \\ -i & 0 \end{pmatrix} \quad , \quad (\sigma^3_{ab}) = (-\sigma_{3ab}) = \sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \end{aligned} \quad (7)$$

where σ_j , $j = x, y, z$ are the known Cayley-Sylvester-Pauli matrices [4, 6, 12-14].

The expressions (5) show explicitly a direct relationship between \underline{L} and \underline{U} , which may be useful in applications of spinorial calculus [11] in Minkowski spacetime.

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